## 8.1 An Aladin-Based Search for Proper-Motion Companions to Young Stars in the Local Association, Tucana-Horologium and $\beta$ Pictoris

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#### Abstract

We have used the Aladin sky atlas of the Virtual Observatory to look for new common propermotion pairs in three young stellar kinematic groups: Local Association ( $\tau \sim 10 - 120$  Myr), Tucana-Horologium ( $\tau \sim 30$  Myr) and  $\beta$  Pictoris ( $\tau \sim 12$  Myr). We have found 9 new and 14 known common proper-motion companions to the 210 investigated stars. With the CAFOS instrument at the 2.2 m Calar Alto telescope, we have investigated in detail one of the new pairs, the HD 143809 AB system, which is formed by a bright G0V primary star and a previously unknown young M1.0–1.5Ve star.

## Introduction

Young nearby late-type stars are excellent targets for high-contrast imaging surveys for brown dwarf and planetary companions. Many of these young low-mass stars in the solar neighbourhood belong to stellar kinematics groups with ages younger than the Pleiades, such as the Local Association and its kinematic subgroups (Montes et al. [2001b]; Song et al. [2003]; Zuckerman & Song [2004]). One way of identifying such stars is searching for faint proper-motion companions at wide separations to already-known members in young stellar kinematic groups.

We followed the procedure described by Caballero et al. [2010a] and used a powerful Virtual Observatory tool, the Aladin sky atlas (Bonnarel et al. [2000]), to look for proper-motion companions to stars in the Local Association (LA,  $\tau \sim 10 - 120$  Myr), Tucana-Horologium (Tuc-Hor,  $\tau \sim 30$  Myr) and  $\beta$  Pictoris ( $\beta$  Pic,  $\tau \sim 12$  Myr) stellar kinematic groups.

### Analysis

We searched for either primary (i.e., brighter) and secondary (i.e., fainter) companions to 210 nearby young stars in the three different moving groups compiled by Montes et al. [2001b] and Torres et al. [2008]: 116 in LA, 44 in Tuc-Hor, and 50 in  $\beta$  Pic. We used the interactive software Aladin v5 to load 2MASS (Skrutskie et al. [2006]) and USNO-B1 (Monet et al. [2003]) astro-photometric catalogues and cross-matched them in a circular area of radius 30 arcmin centred on each target. Next, we constructed a proper-motion diagram with the Aladin application VOPlot and searched for sources with USNO-B1 proper motions different from those of the target stars by less than 10 mas/yr.

For each proper-motion candidate, we searched the literature for previous claims of membership in multiple systems (e.g., the Washington Double Star catalogue; Mason et al. [2001]) and for better proper-motion determinations (e.g., Tycho-2; Høg et al. [2000]). We derived photometric distances for both primaries and secondaries, based on available spectroscopy or photometry, of the candidate pairs without parallactic distance measurements and discarded those with no coincident values.

### Results

Of the 210 investigated stars, we identified 23 multiple system candidates, of which 14 were known common proper-motion companions and 9 were unknown multiple systems (see Table 8.1 and Fig. 8.1).

Name	Sp.	ρ	$\theta$	s	SKG
	$\operatorname{type}$	[arcsec]	[deg]	[kAU]	
HD 82939	G5V	$162.28 \pm 0.17$	$121.49 \pm 0.07$	$6.3\pm0.2$	LA
GJ 9303	K7V				
EX Cet	G5V	$612.10\pm0.11$	$258.66\pm0.02$	$14.7\pm0.3$	LA
G 271–110	M3.5V				
HD 143809 A	G0V	$86.40\pm0.11$	$252.57\pm0.09$	$7.1\pm0.9$	LA
HD 143809 B	M1.0-1.5V				
HD 13183	G7V	$705.99 \pm 0.10$	$103.761\pm0.009$	$36.0\pm1.2$	Tuc-Hor
CD-53 413	G5V				
CD $-53~544$	K6Ve	$22.06\pm0.08$	$11.11\pm0.17$	$0.93\pm0.09$	Tuc-Hor
AF Hor	M2Ve				
HD 207964 AB $$	F1III+	$1412.75\pm0.11$	$245.020\pm0.005$	$64.0\pm1.9$	Tuc-Hor
HD 207575	F6V				
HD 173167	F5V	$550.31 \pm 0.10$	$290.244\pm0.011$	$29 \pm 3$	$\beta \operatorname{Pic}$
TYC 9073–0762–1	M1Ve				
$\eta$ Tel AB	A0Vn+M7.5V	$416.26\pm0.13$	$170.691\pm0.012$	$20.1\pm0.2$	$\beta \operatorname{Pic}$
HD 181327	F6V				
HD 199143 AB	F7V+	$325.04 \pm 0.08$	$138.35\pm0.02$	$14.8\pm0.5$	$\beta \operatorname{Pic}$
AZ Cap	K6Ve				

Table 8.1: The nine unknown proper-motion pairs (SKG indicates the stellar kinematic group).

One of nine of them was a suspected multiple system ( $\eta$  Tel AB and HD 181327; Schneider et al. [2006]). At the measured or derived distances, the angular separations of 0.37 to 24 arcmin translate into projected physical separations between 0.0045 and 0.31 pc. Interestingly, some of the proper-motion companions had already been tabulated as members in the same stellar kinematic group as their target stars (e.g., HD 207964 AB and HD 207575 in Tucana-Horologium).

One of the new multiple systems was formed by a young solar analogue and an anonymous high proper-motion red dwarf never described in the literature before, and was subject of a dedicated astrometric, photometric and spectroscopic follow-up study. First, we confirmed the common proper motion of the "HD 143809 AB" system using 11 astrometric epochs separated by over 56 years as in Caballero et al. [2010b]. Next, we collected B, V, R and I images and low-resolution spectra (grating G100) with the CAFOS instrument at the 2.2 m Calar Alto telescope. While the primary is a known G0V star with a high lithium abundance (EW(Li I) = 103 mÅ; López-Santiago et al. [2010]), an estimated age of  $\tau \sim 80 - 120 \text{ Myr}$  and kinematics consistent with membership in the Local Association Montes et al. [2001a], the new companion at  $s = 7.1 \pm 0.9 \text{ kAU}$  is an M1.0–1.5Ve star with chromospheric H $\alpha$ , H $\beta$  and H $\gamma$  emission. The heliocentric distance derived from its spectral type and photometry matches the one of the primary measured by *Hipparcos* at  $d = 78 \pm 8 \text{ pc}$ . Using this distance, its *J*-band apparent magnitude ( $J = 10.35 \pm 0.03 \text{ mag}$ ) and the NextGen models (Baraffe et al. [1998]) for an age of 100 Myr, HD 143809 B has a most probable mass of  $0.57 - 0.60 \text{ M}_{\odot}$ .

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Figure 8.1: False-colour images combining DSS POSSII  $B_J$ ,  $R_F$  and  $I_N$  photographic plates of the nine new pairs. Labelled are the multiple system names and field-of-view sizes. North is up and east is to the left. The red arrows show the proper motion.

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